

Towards an ECAL e^- analysis paper?

David Ward

- ❖ Would like to aim to progress from LCWS note to published paper(s). Envisage up to 4 papers:
 - ❖ 1. Hardware, calibration, technical performance (gain, noise, stability etc). Anne-Marie to coordinate.
 - ❖ 2-4. Electron response
- ❖ Schedule – drafts by the end of 2007???
- ❖ Basic topics much as in the LCWS note.
- ❖ Several problems to be addressed before we can be ready. **Will mainly discuss these.**
- ❖ Comments on systematics.

Content of analysis papers?

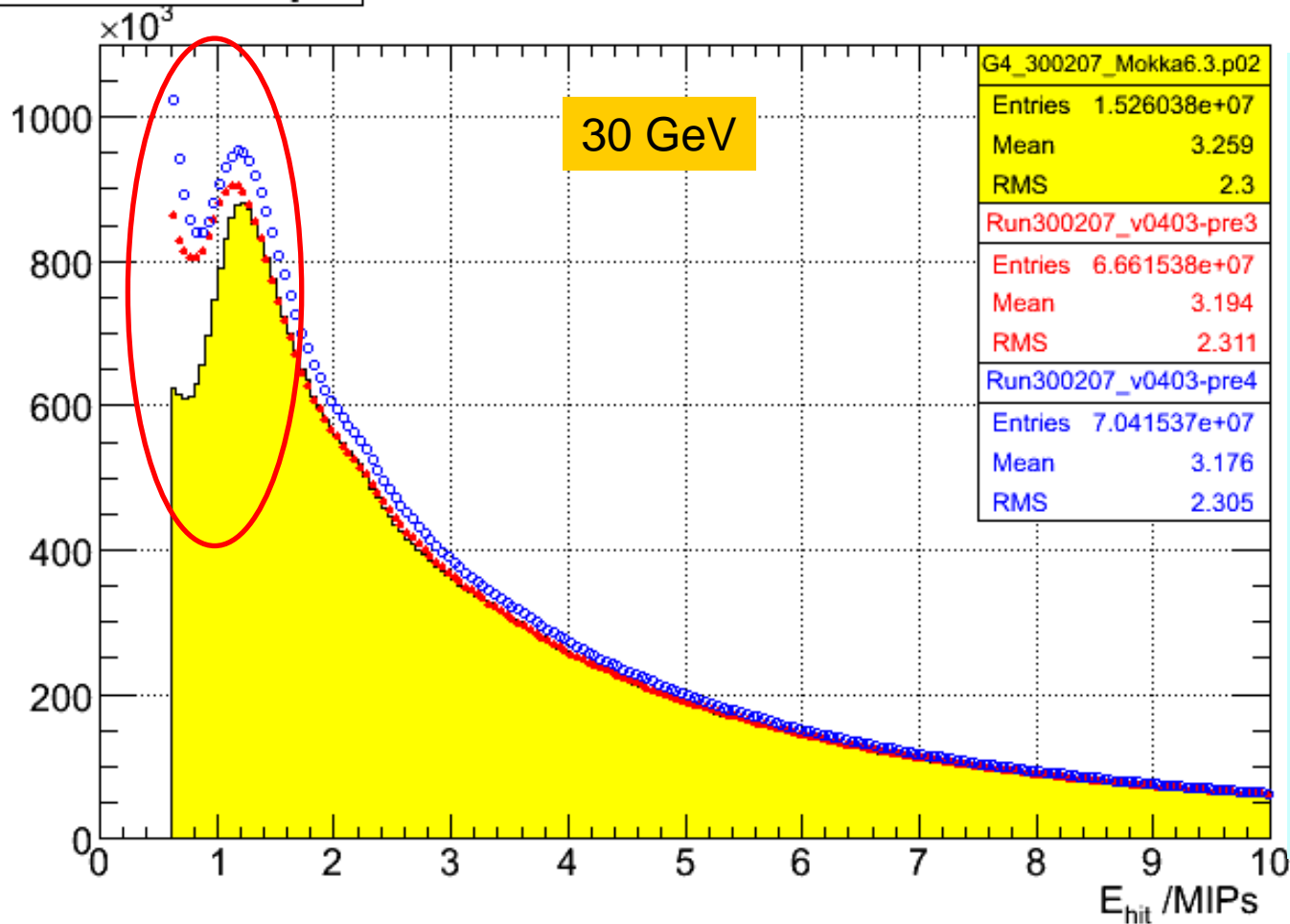
- C.Carloganu, DRW...
 - ❖ Energy response
 - ❖ Linearity; uniformity (gap correction); angle dependence.
 - ❖ Energy resolution
 - ❖ As a function of energy, angle, impact of gaps...
- Valeria, George, Nige
 - ❖ Shower longitudinal profile
 - ❖ Shower transverse profile
 - ❖ Effective Molière radius
- Paul, Hakan, Michele
 - ❖ Position resolution
 - ❖ Using tracking
 - ❖ Angular resolution
 - ❖ Using tracking
 - ❖ 2-shower separation
 - ❖ Using double events/superimposed events
 - ❖ All of the above compared with MC simulation.

ECAL problem areas

- ❖ There are several problematic areas, where we see discrepancies between data and Monte Carlo.
- ❖ Some at least may be partly interconnected.
 - ❖ Low pulse height hits
 - ❖ Number of hits.
 - ❖ Interwafer gaps (effect of guard rings etc)
 - ❖ Transverse shower shape ($\sim 10\%$ higher in data)
 - ❖ Shower depth (understanding of beam line, upstream material?)
 - ❖ Mismatch of energy scale ($\sim 3\%$) between CERN and DESY

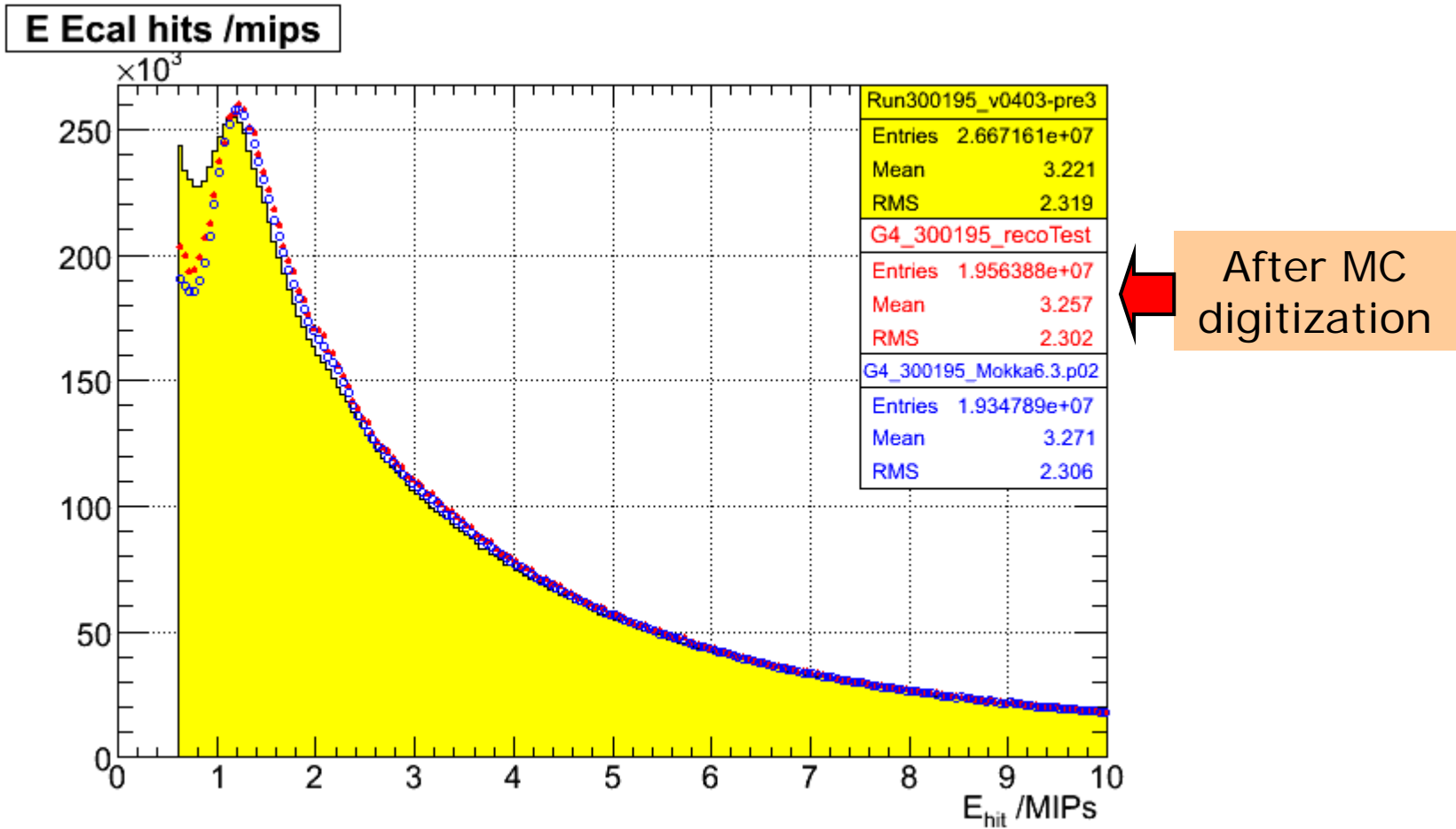
Low energy hits

E Ecal hits /mips



- ❖ Pre3 processing agrees with MC quite well down to 1.2 MIPs
- ❖ Pre4 processing (SIPS correction) gives extra hits; agrees less well with MC.
- ❖ Effect grows with increasing energy.
- ❖ Anne-Marie's MC digitization and reconstruction doesn't have any significant effect (next slide...).
- ❖ Similar in 2007 data; possibly a bit worse

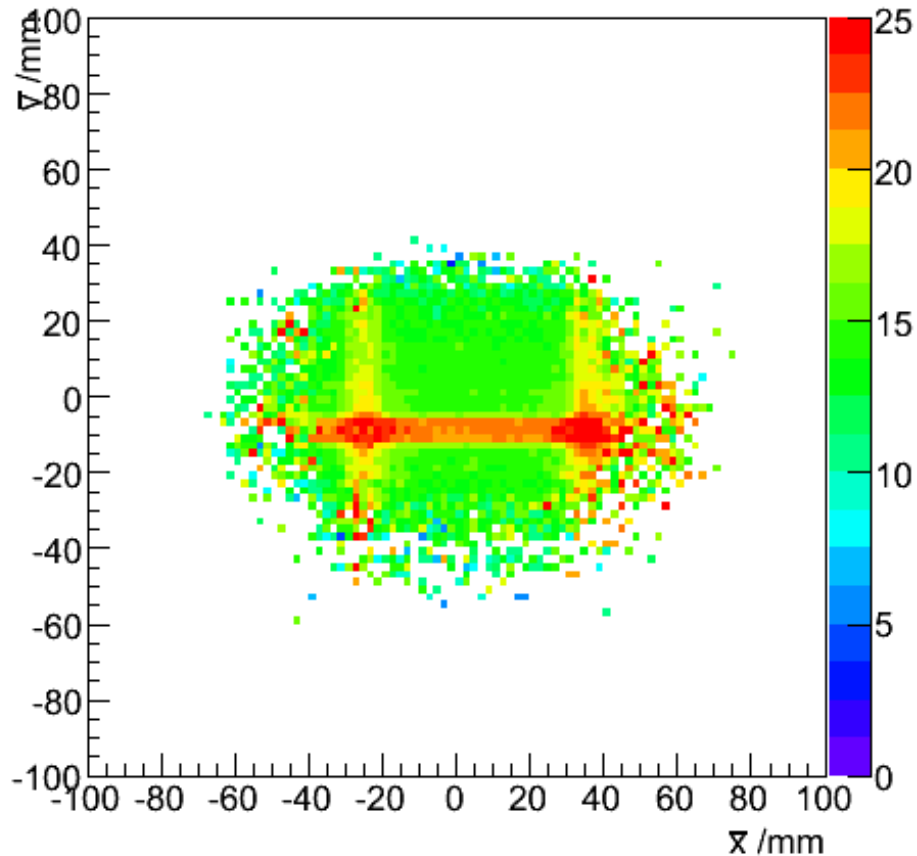
Effect of MC digitization (45 GeV)



Correlated with shower position

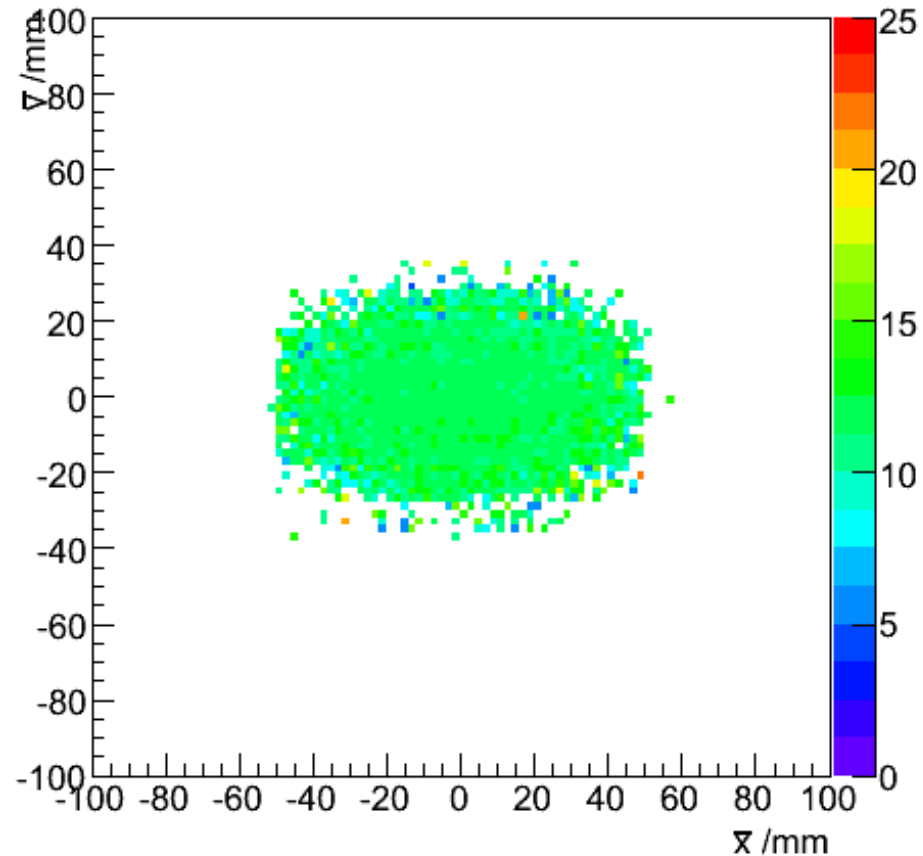
$N(<0.8\text{MIP})$

2006 Data 30 GeV



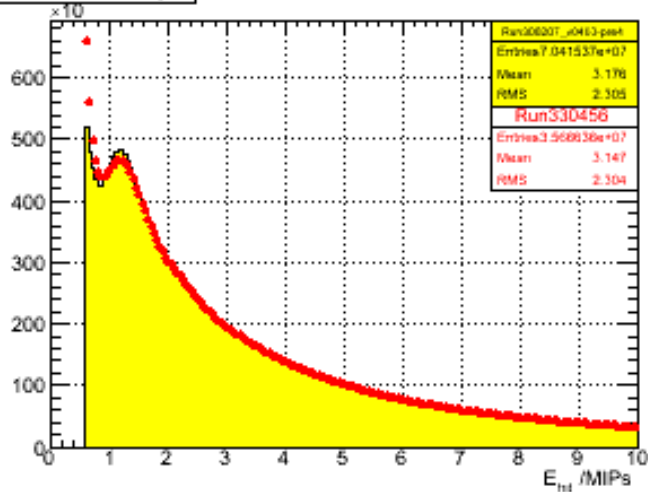
$N(<0.8\text{MIP})$

MC 30 GeV

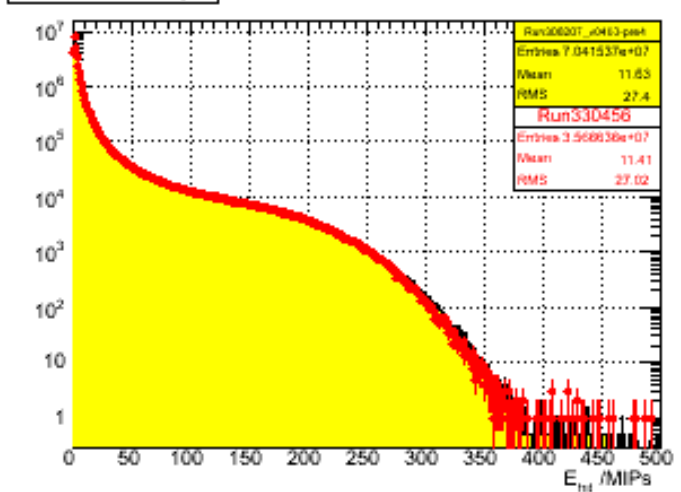


2006 *c.f.* 2007 – 30 GeV e^-

E Ecal hits /mips

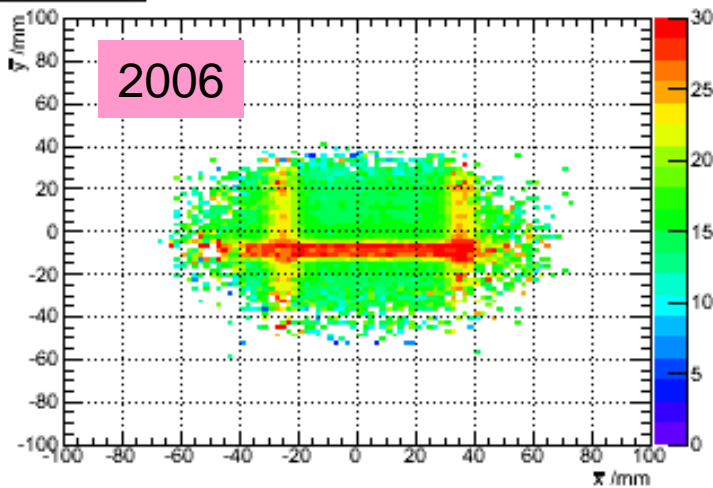


E Ecal hits /mips

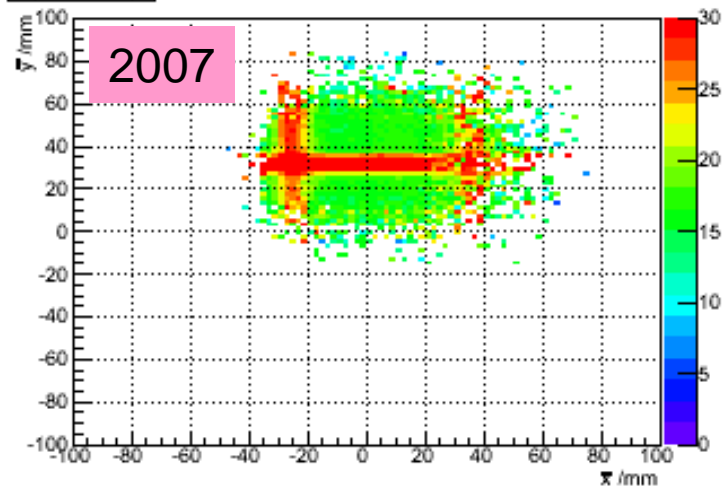


A little worse in 2007?

N(<0.8MIP)



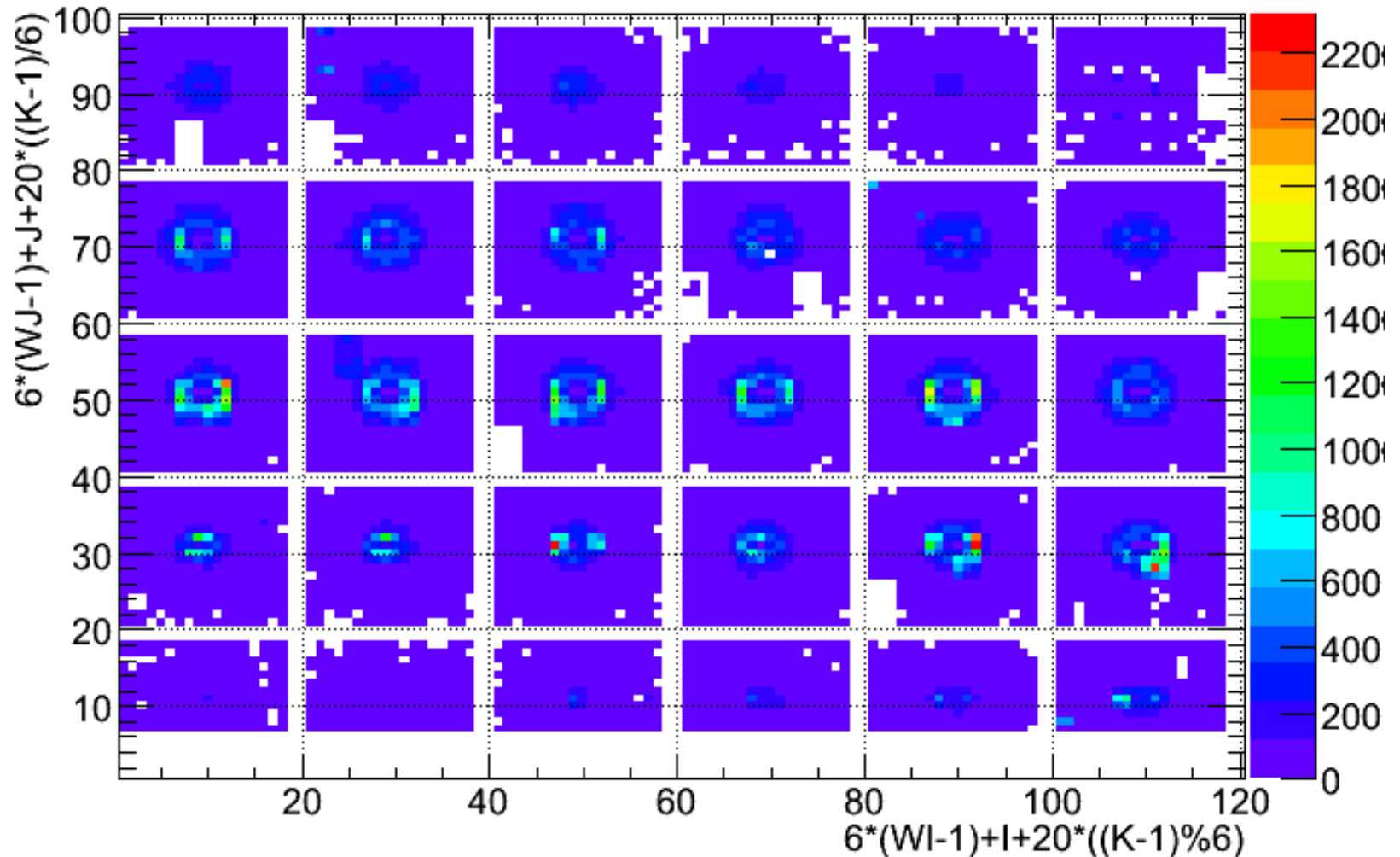
N(<0.8MIP)



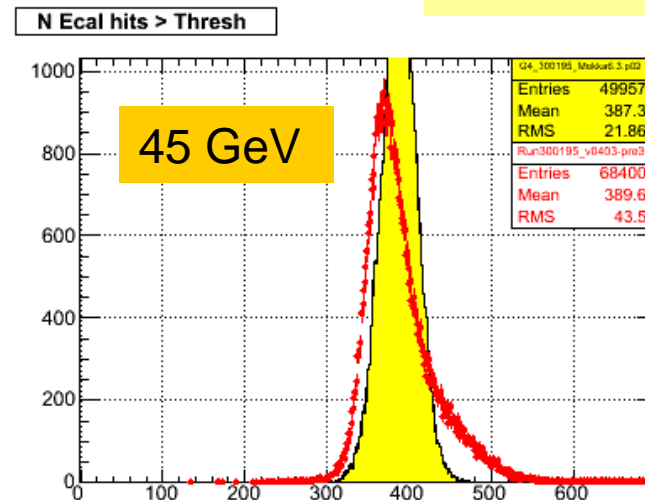
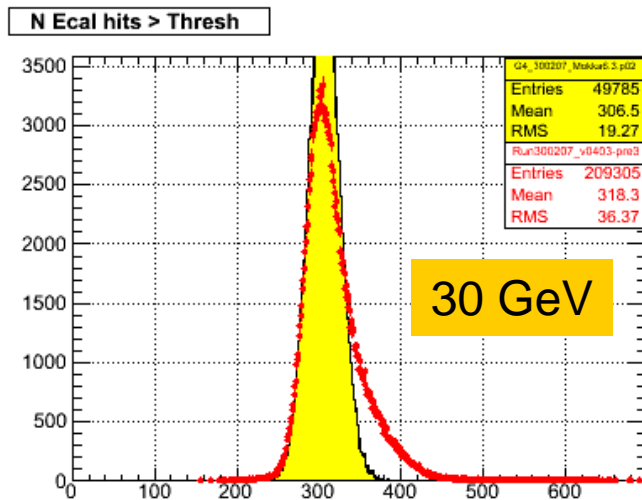
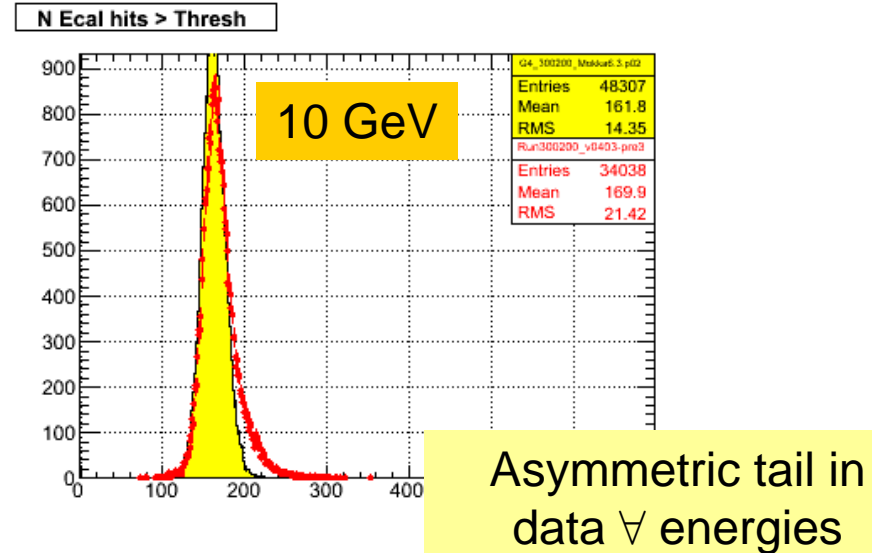
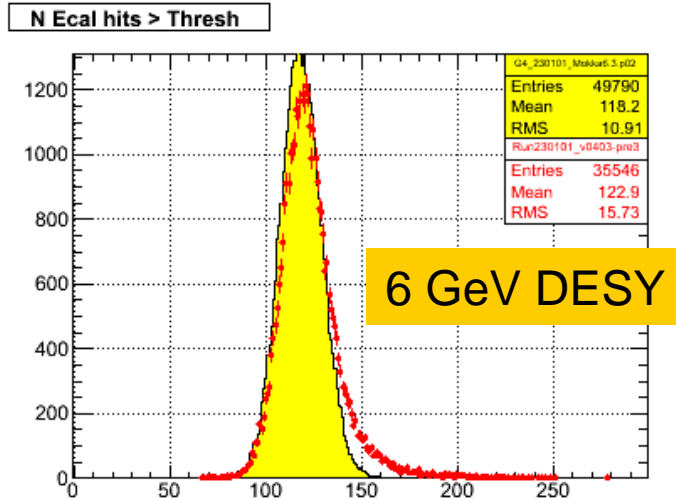
HitMap of hits below 0.8 MIP

HitMap ECAL Ehit<0.8MIP

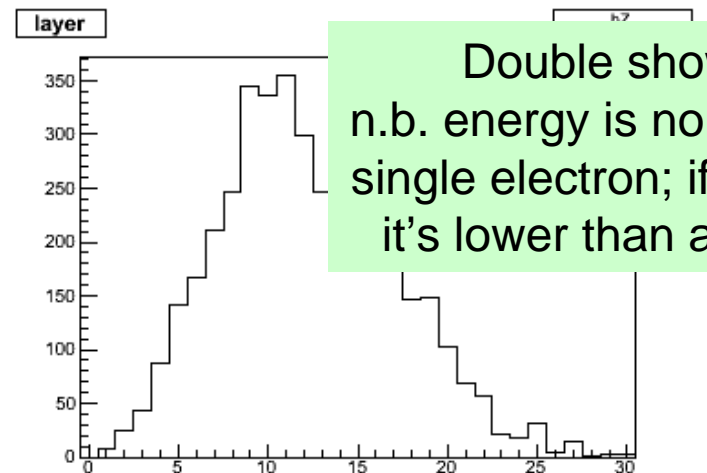
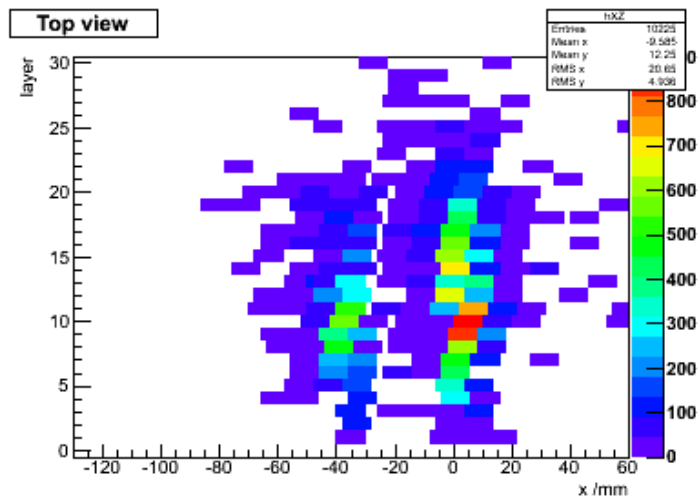
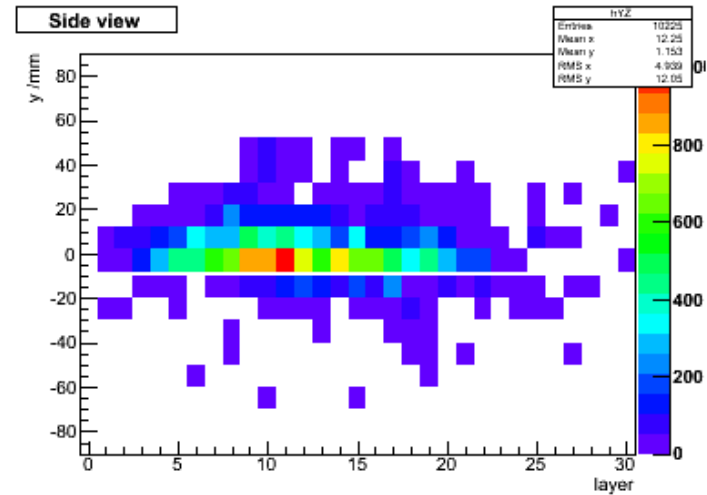
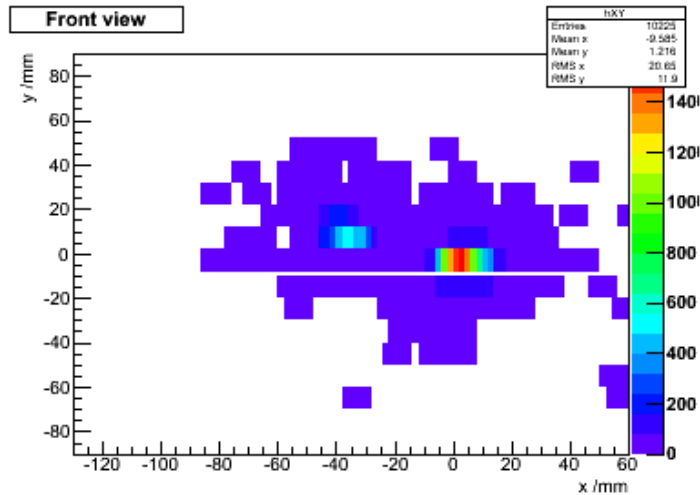
2007 Data



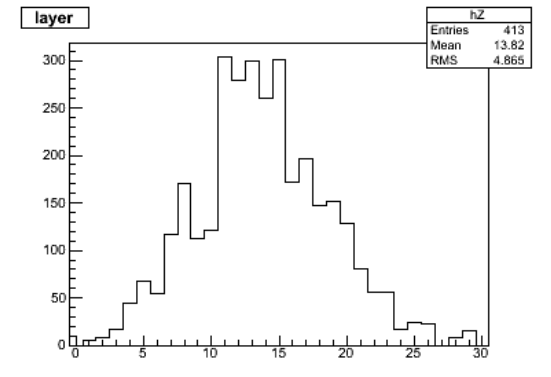
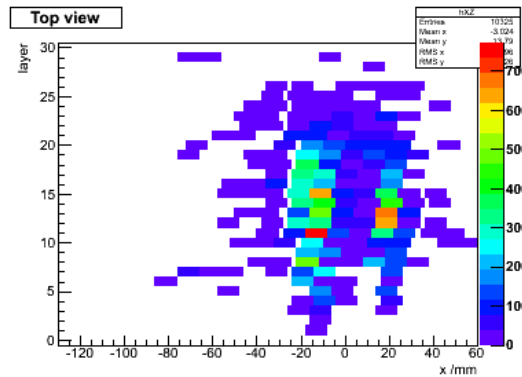
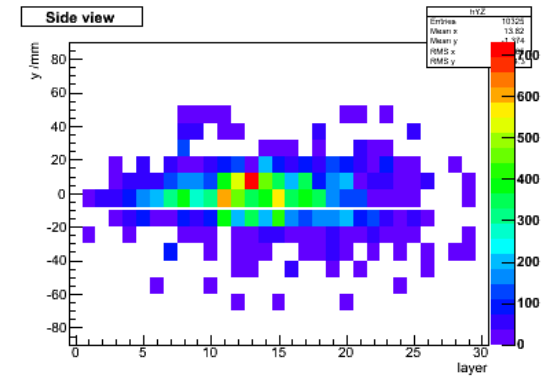
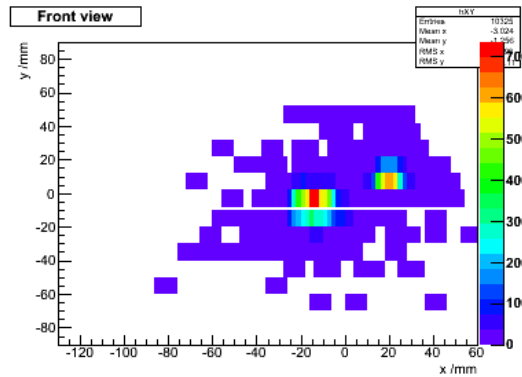
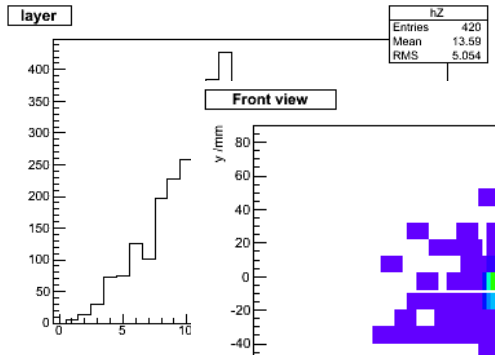
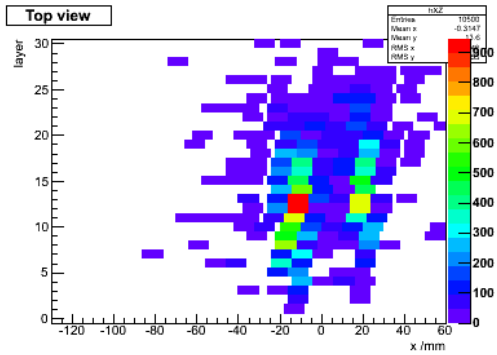
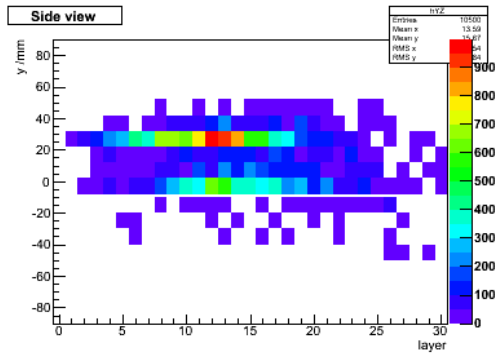
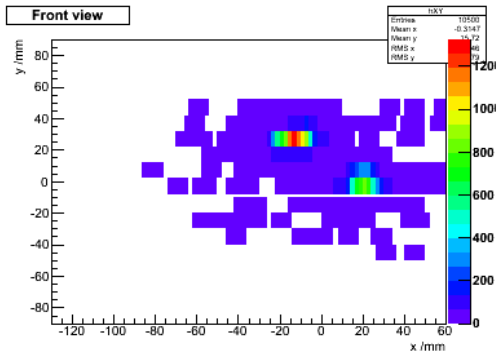
Number of hits (above threshold)



Look at events in the high tail in data



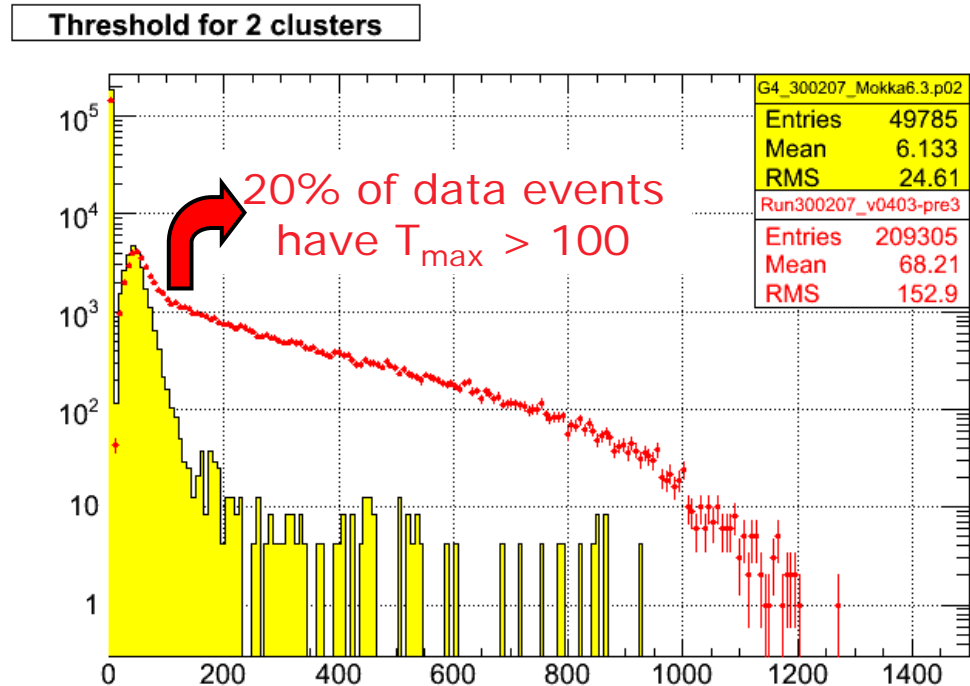
Double shower.
n.b. energy is normal for a
single electron; if anything
it's lower than average.



Such events are essentially not simulated at all in Mokka. There is much still to understand about the beam!

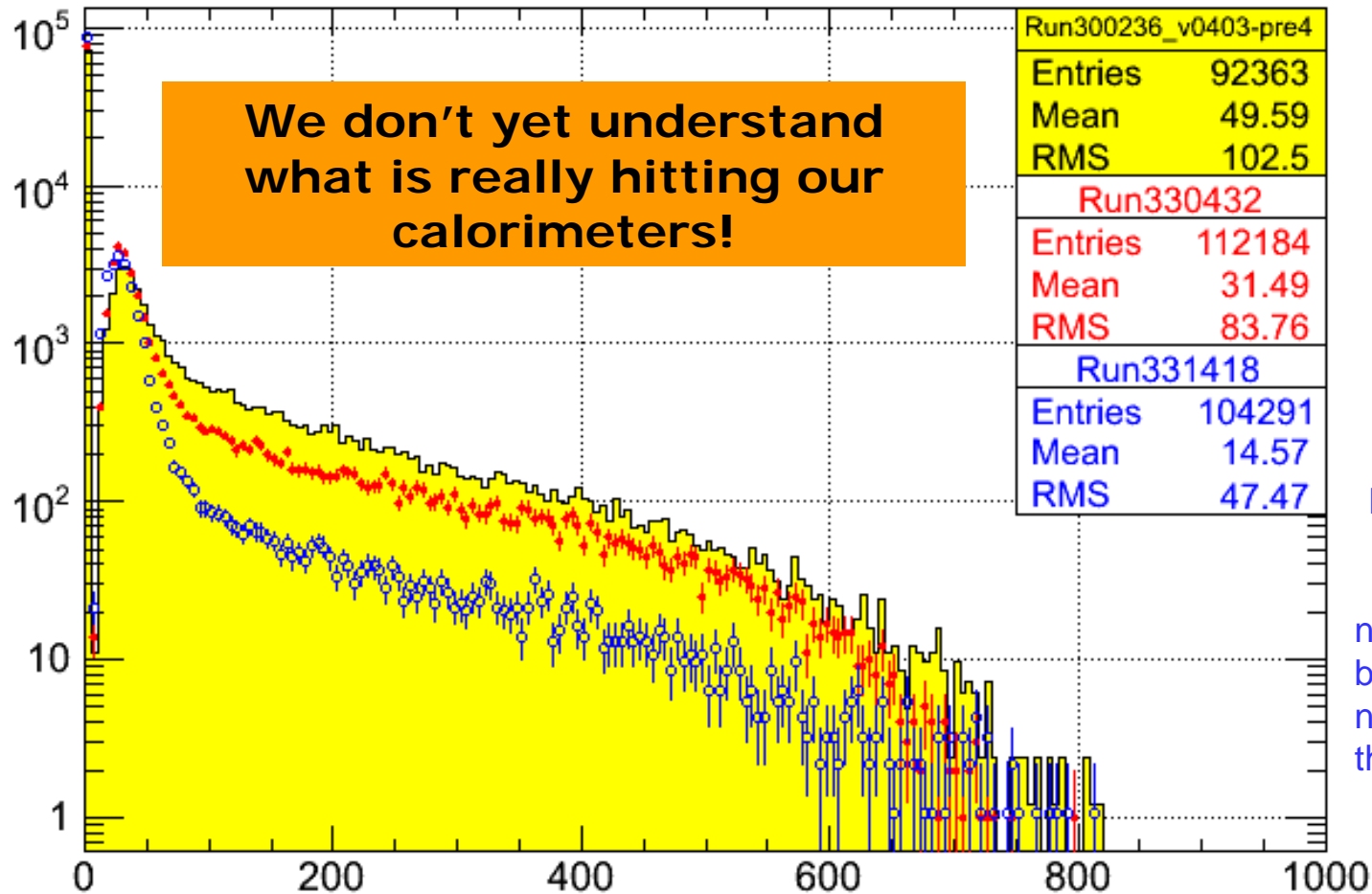
Cut against double events

- ❖ Simple algorithm: form x - y projection of all 30 layers.
- ❖ Apply Threshold T
- ❖ Perform naive nearest-neighbour clustering of cells above threshold.
- ❖ Find T_{\max} below which two or more clusters first appear.
- ❖ Dramatic difference between data and MC
- ❖ Cut $T_{\max} < 100$ to remove (most) double events.



Compare with 2007 data (20 GeV)

Threshold for 2 clusters



2006

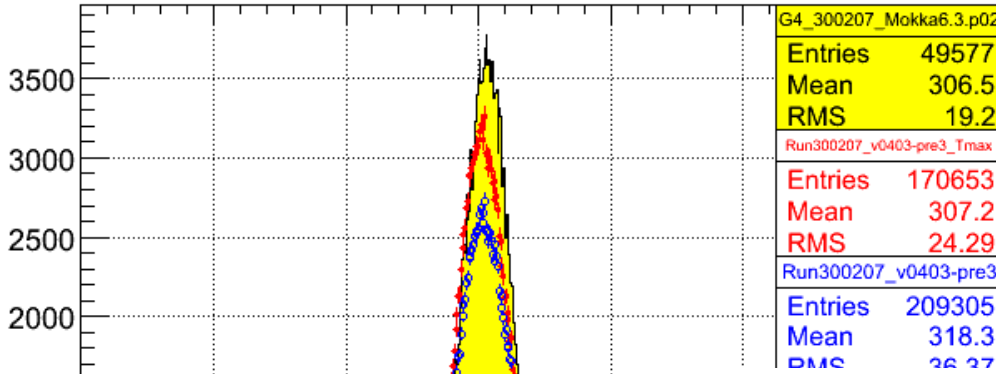
2007 e⁻
Better, but
other runs
worse

2007 e⁺
Much better,
but still worse
than MC

n.b. the e⁺
beam is much
narrower than
the e⁻ beam

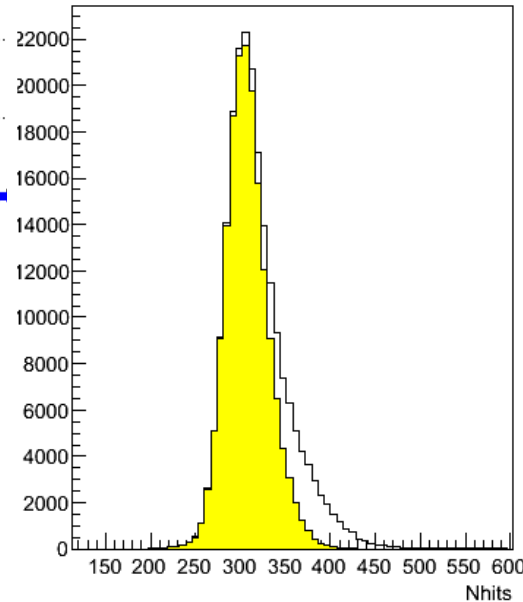
Effect of cut against double showers

N Ecal hits > Thresh

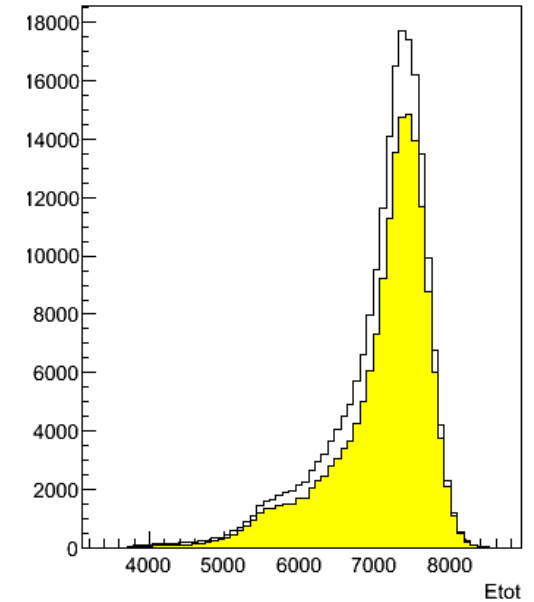


Helps with N_{hits} , but doesn't fully solve the problem.
Little effect on E_{tot} .

Nhits

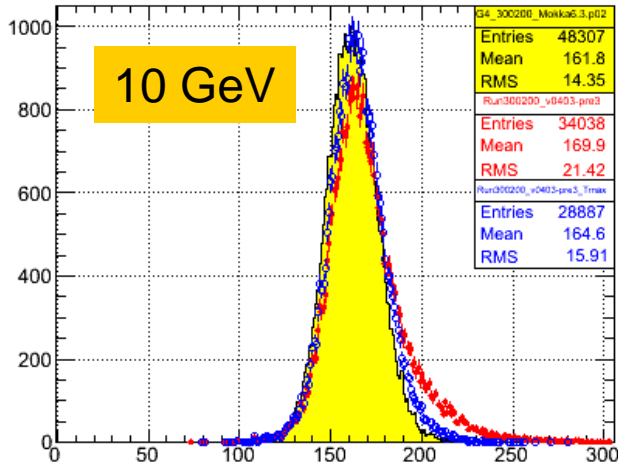


Etot

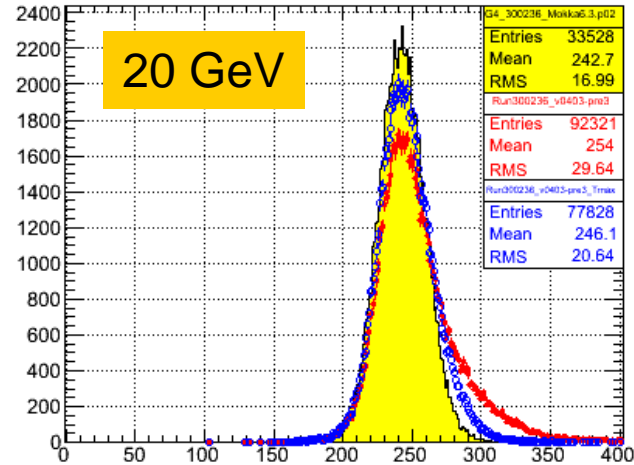


Other energies?

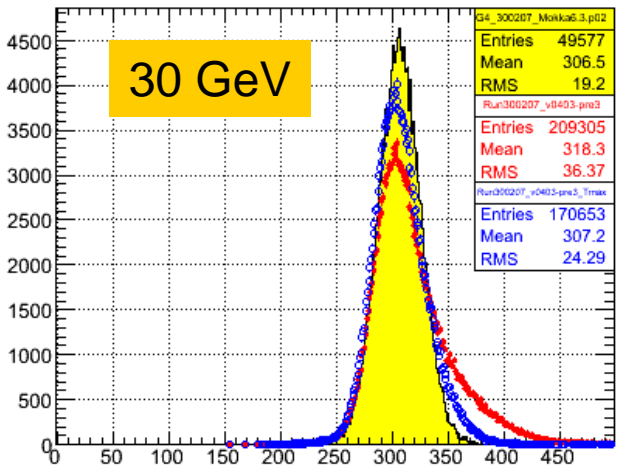
N Ecal hits > Thresh



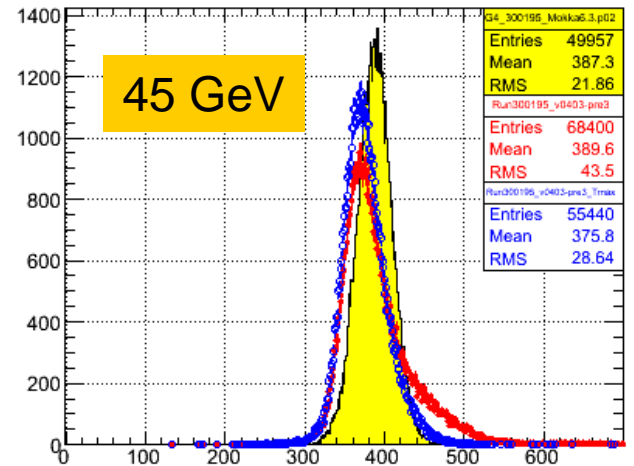
N Ecal hits > Thresh



N Ecal hits > Thresh



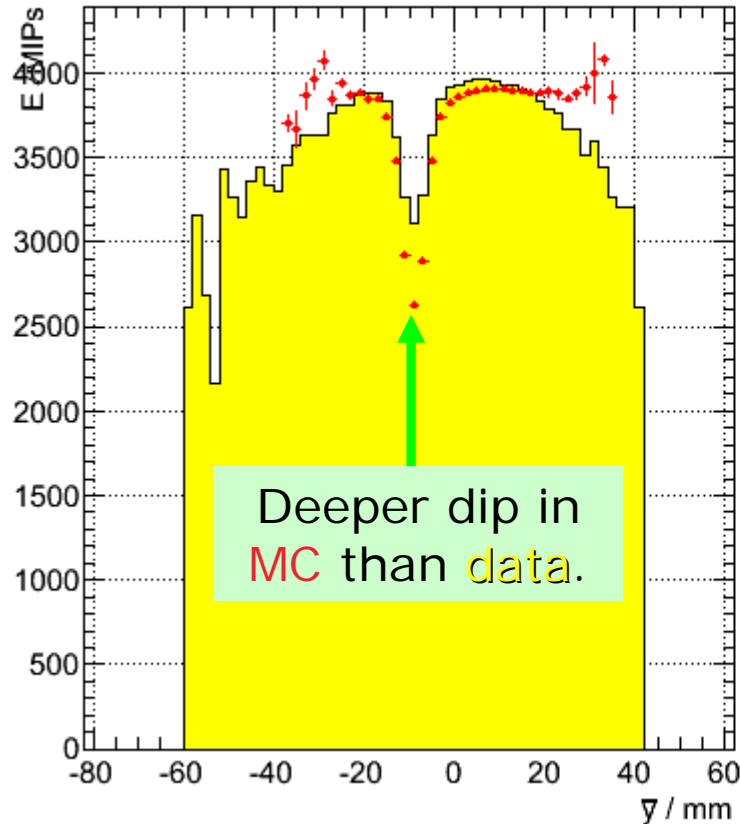
N Ecal hits > Thresh



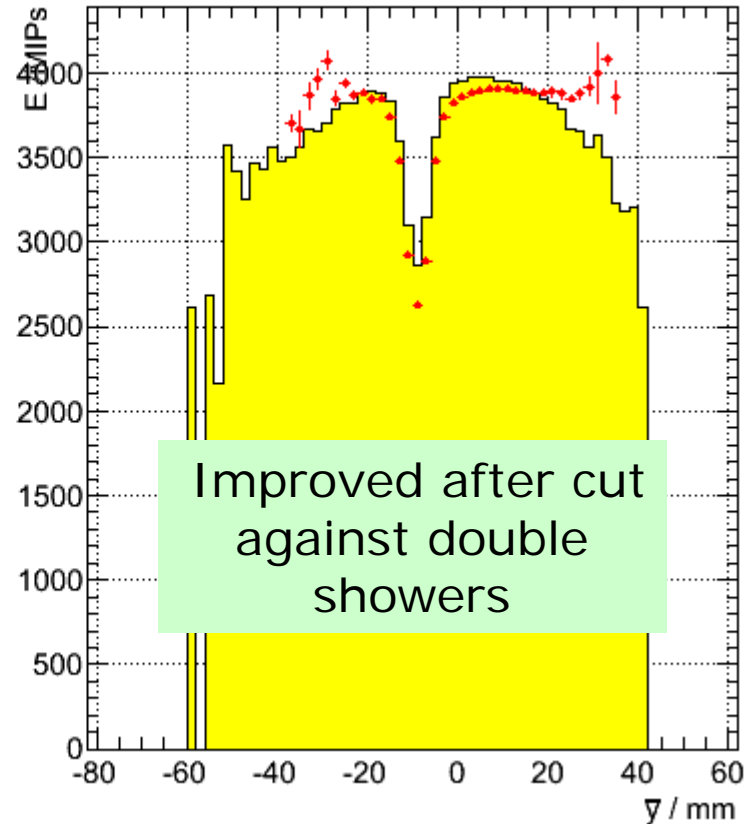
Peak position vs. energy is still a issue

Interwafer gaps; 30 GeV e^-

E vs γ

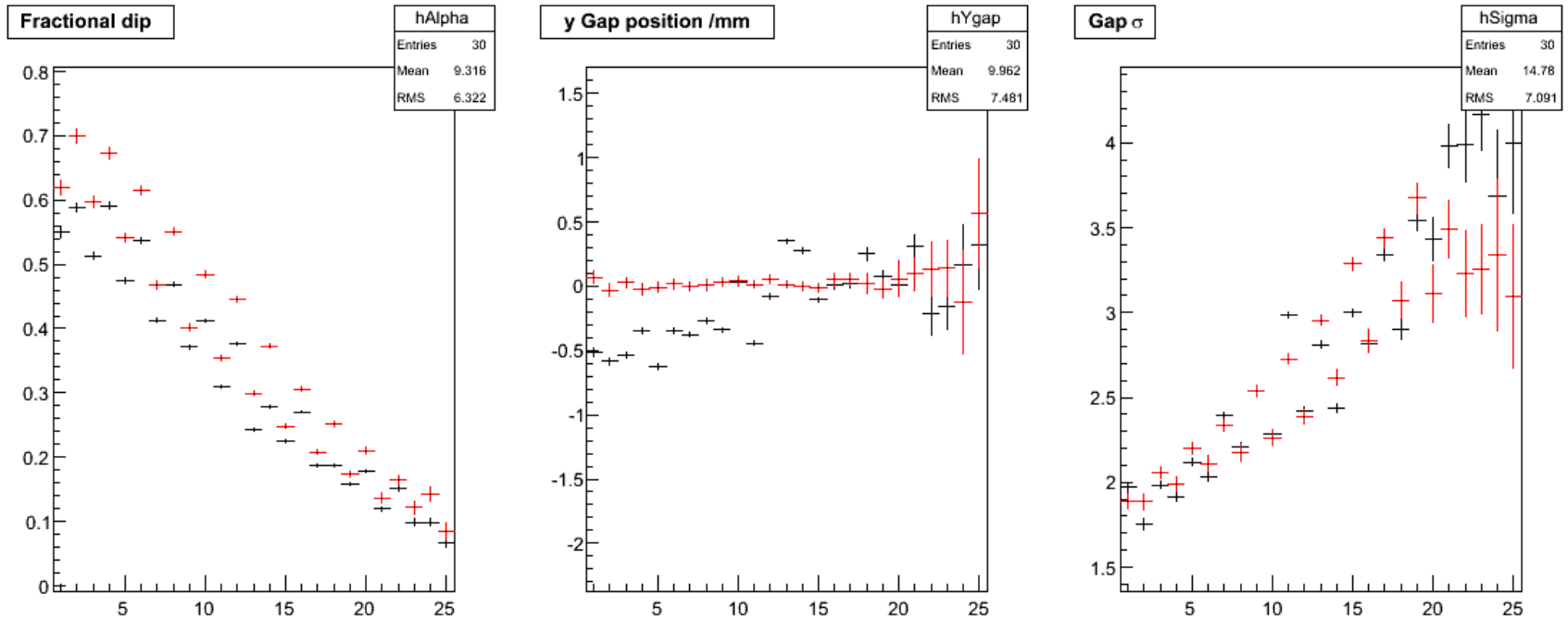


E vs γ



Dip vs. Layer; 30 GeV e^-

Fit dip in energy vs y in each layer to a Gaussian.
Plot fractional depth of dip, position and σ , for data and MC



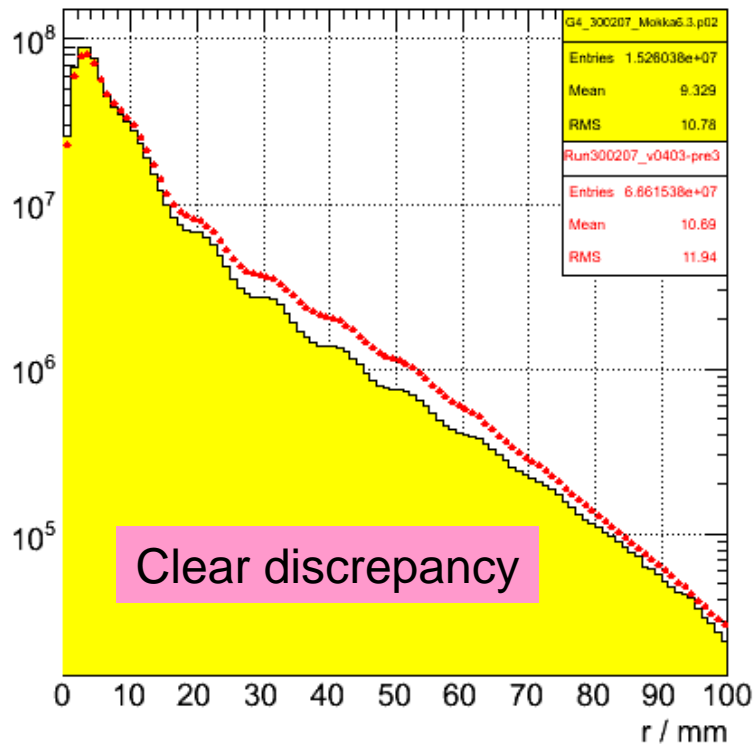
Dip in MC still slightly deeper than in data

Beam inclined in data?
Alignment issues?

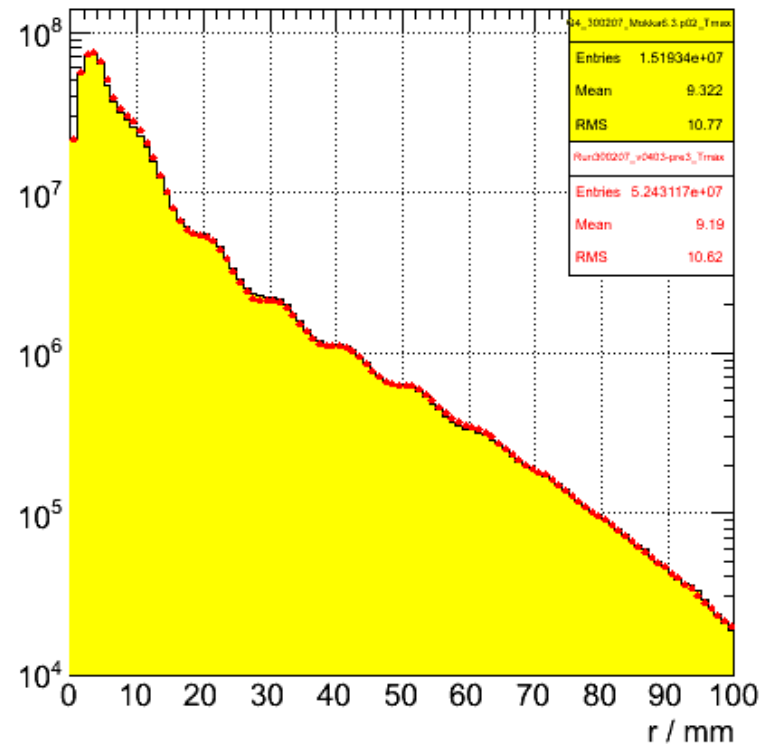
Radial energy distribution

Greatly improved by cut against double showers; at all CERN energies

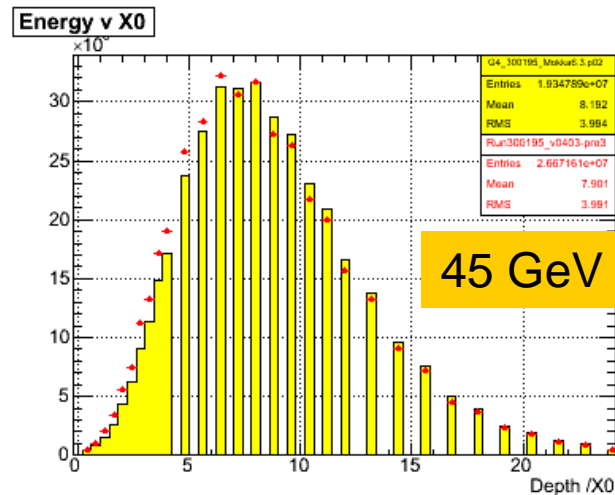
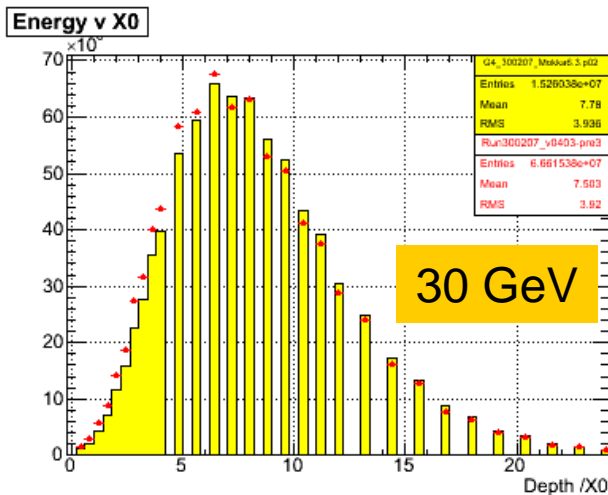
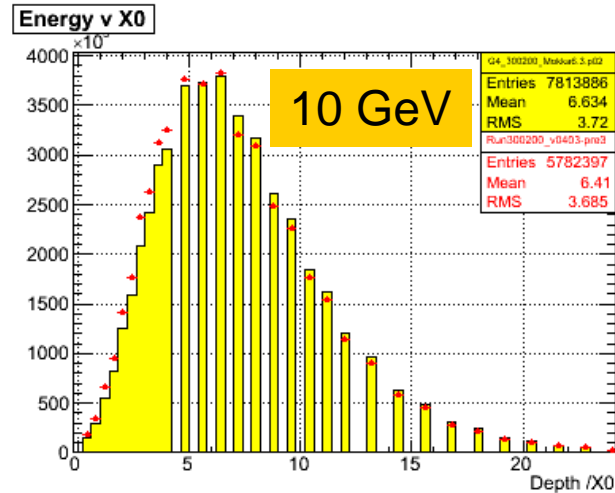
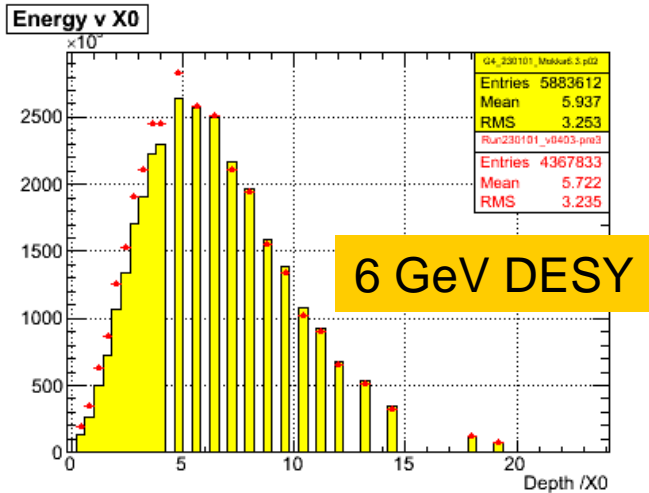
E vs r_{hit}



E vs r_{hit}



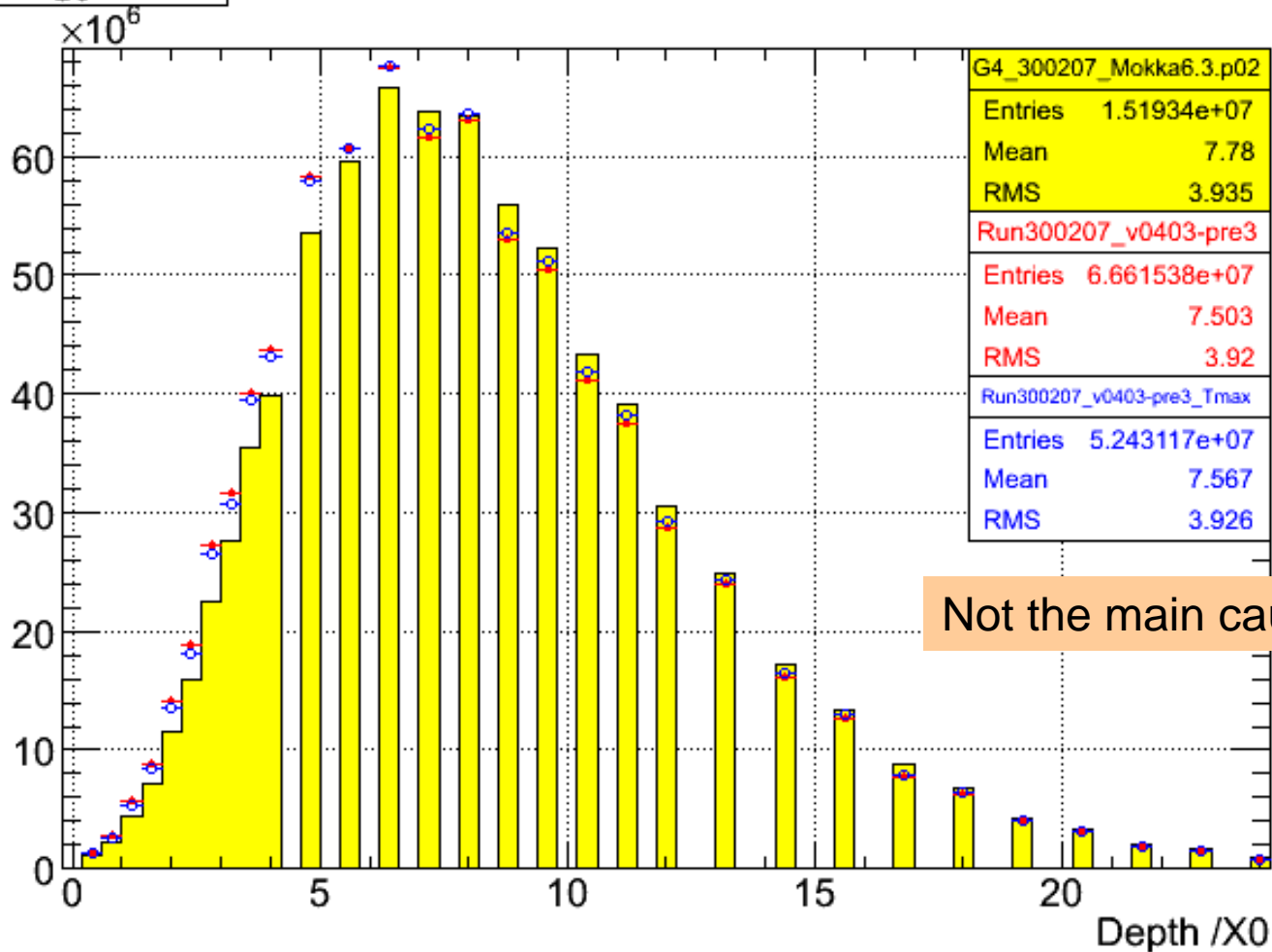
Shower depth



- Shower depth consistently » 0.2-0.3 X_0 deeper in MC than data.
- R.M.S. seems quite well modelled.
- Suggests problem with modelling the beam line rather than the material of ECAL?

Shower depth – effect of double shower cut

Energy v X0



MC 7.78

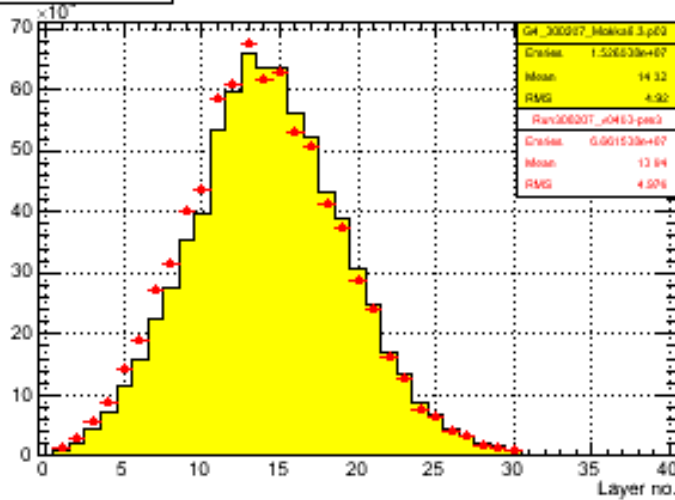
Data 7.50

After T_{max} cut 7.57

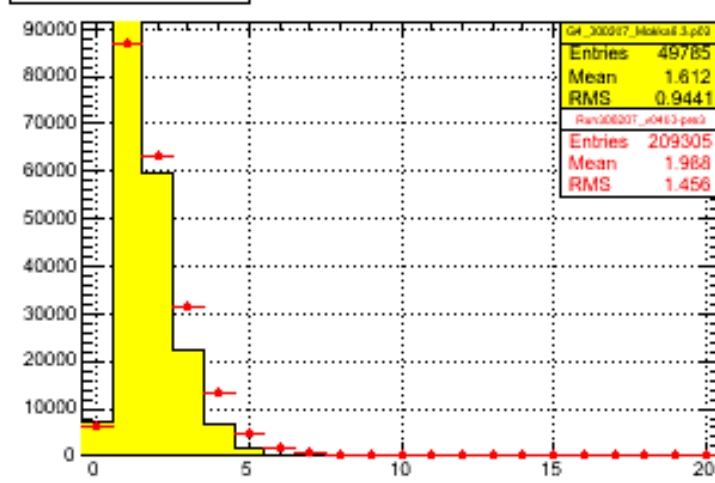
Not the main cause of the problem.

Hits in Layer 1 only (30 GeV e⁻)

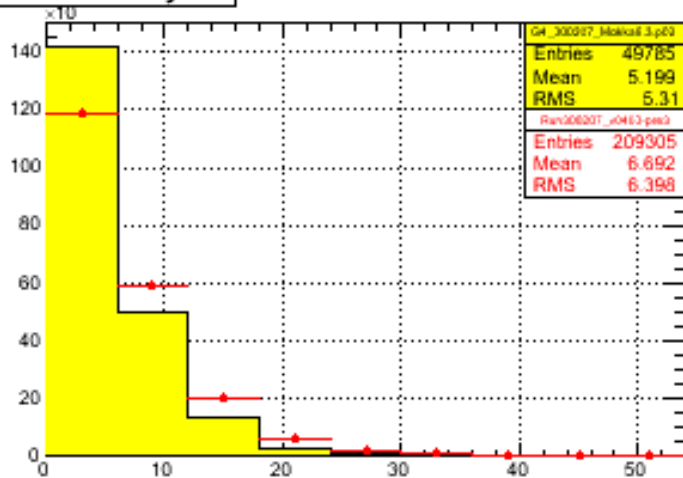
Energy v Plane



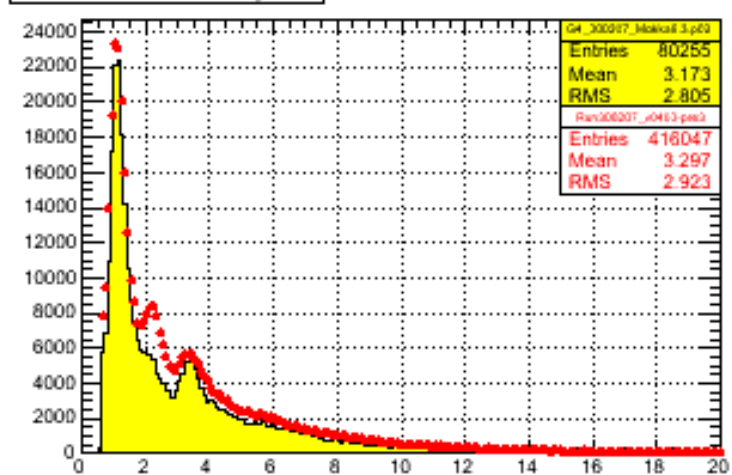
N Ecal hits layer 1



E Ecal /MIPs layer 1

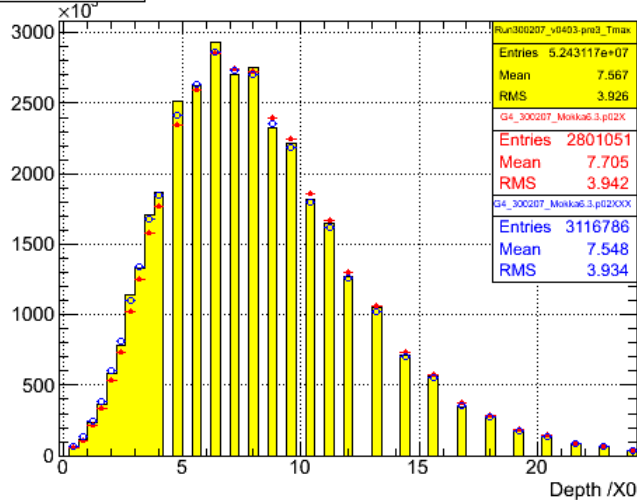


E Ecal hits /MIPs layer 1

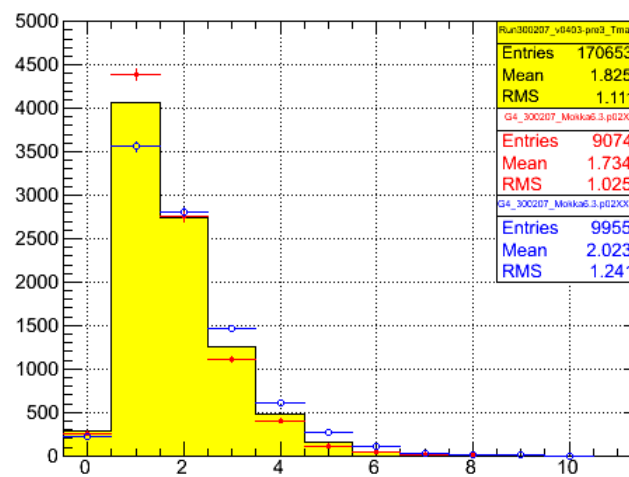


Increase material in beam

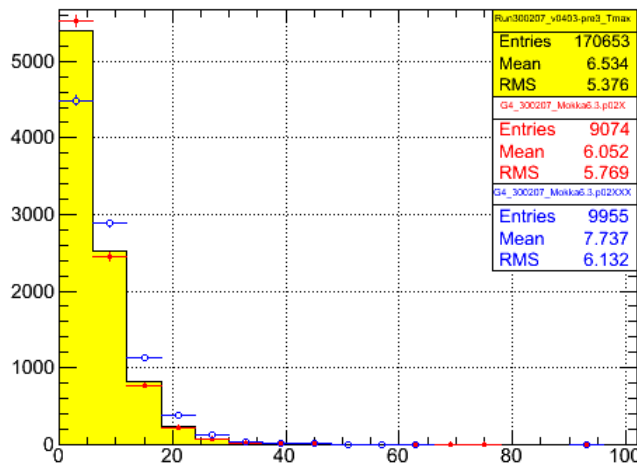
Energy v X0



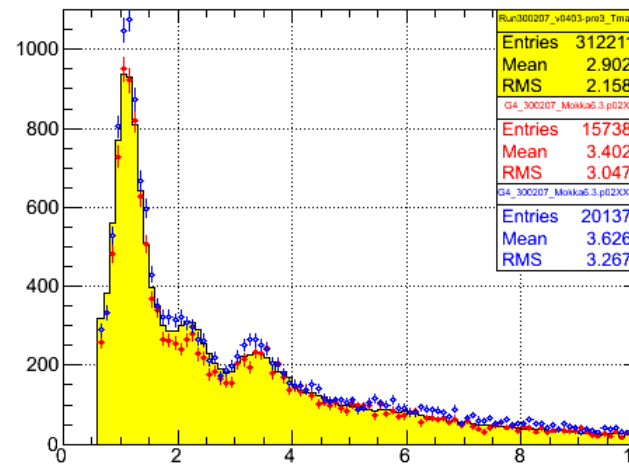
N Ecal hits layer 1



E Ecal /MIPs layer 1

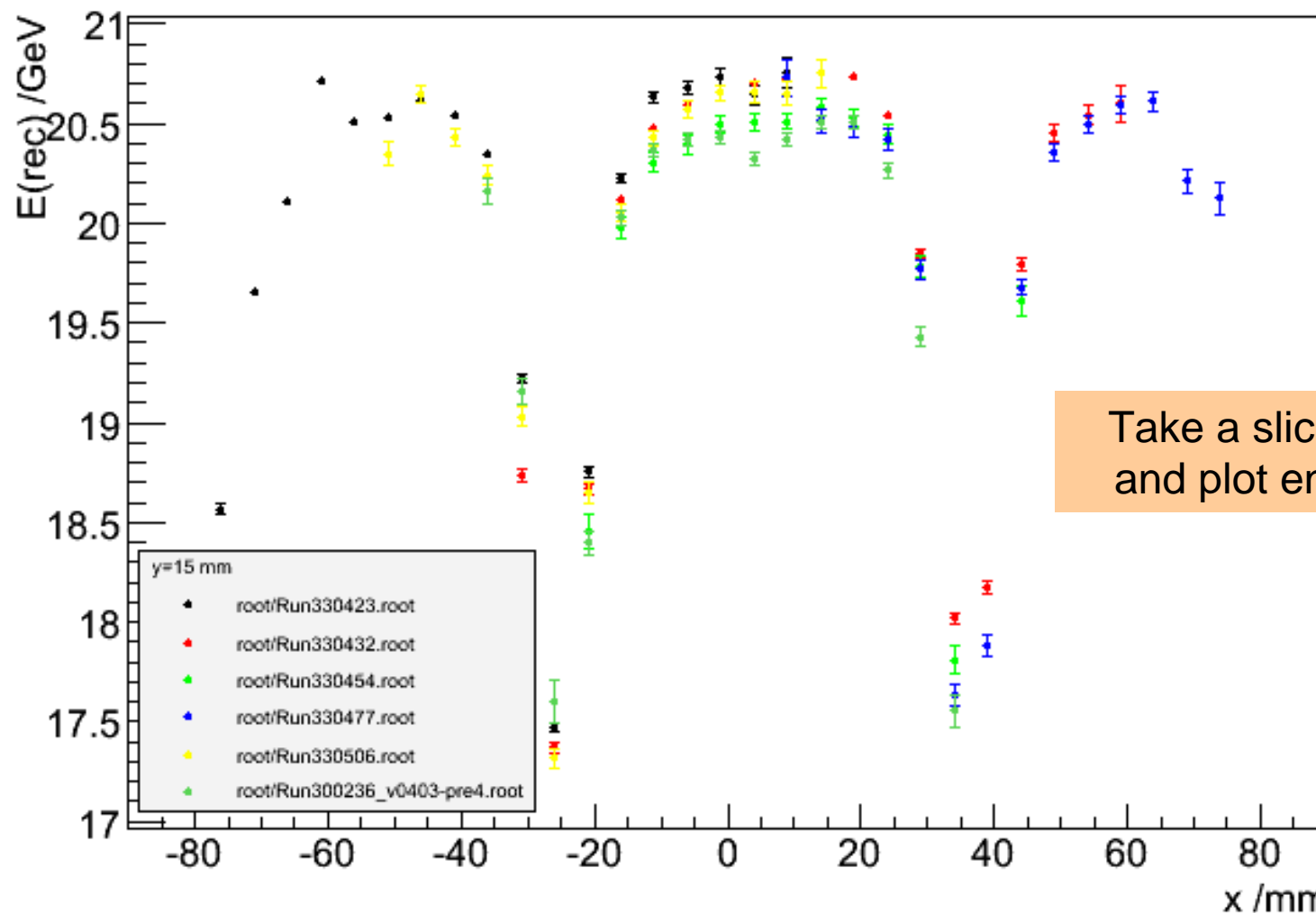


E Ecal hits /MIPs layer 1

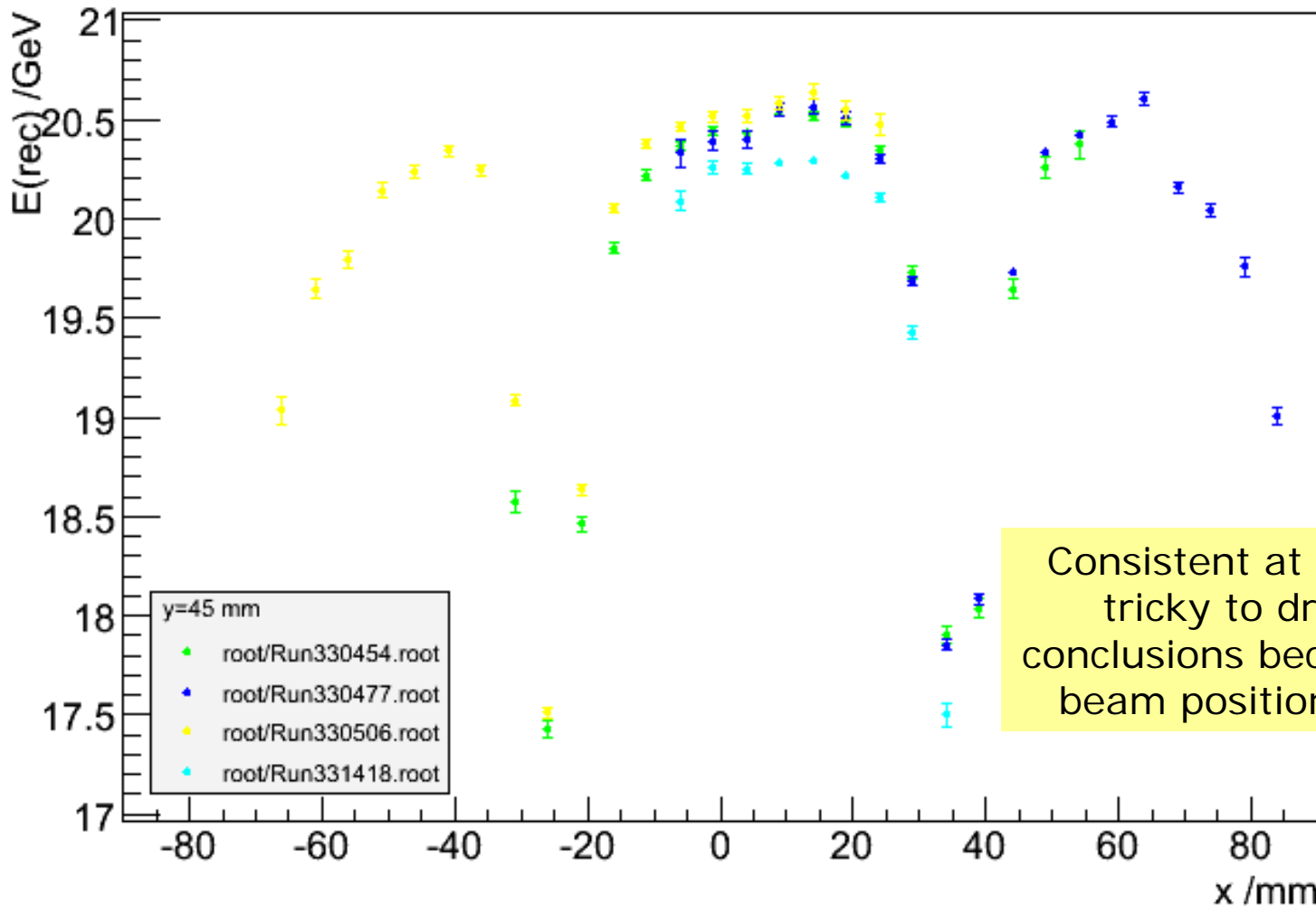


- Try increasing material in beam by 10% or 20% X_0
- Technically done by doubling or trebling thickness of scintillators.
- Obviously too crude, but suggests ~ 10-20% X_0 upstream material would be needed.
- Seems improbably much?
- Emphasises the need to understand the beam better.

Uniformity across detector in 2007 data

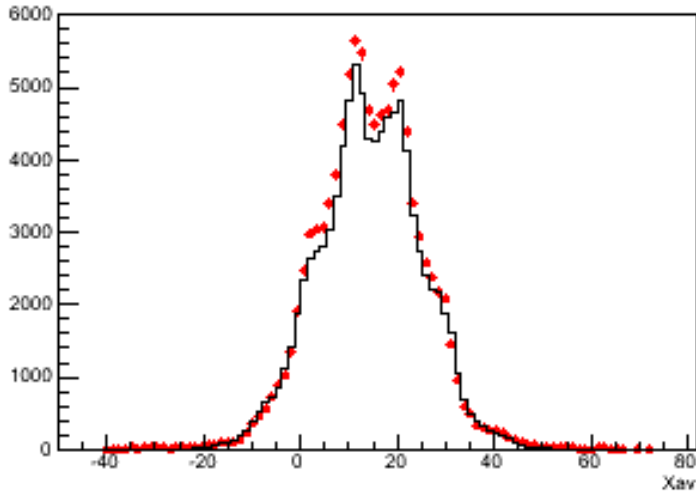


Another slice ($40 < y < 50$)

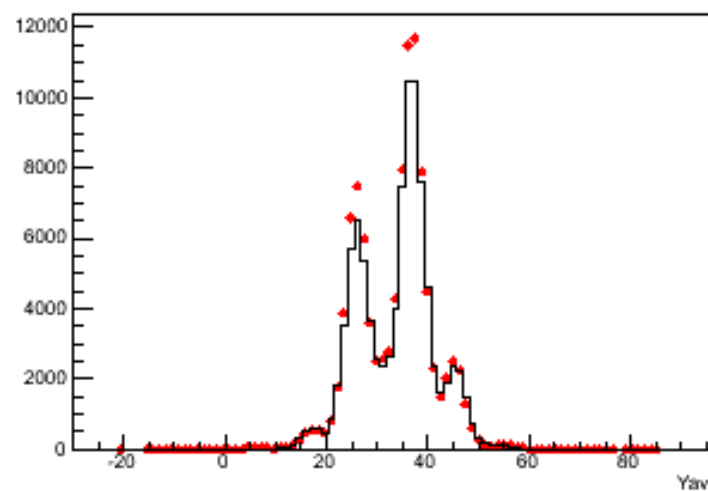


Reweight two 20 GeV runs to same beam profile

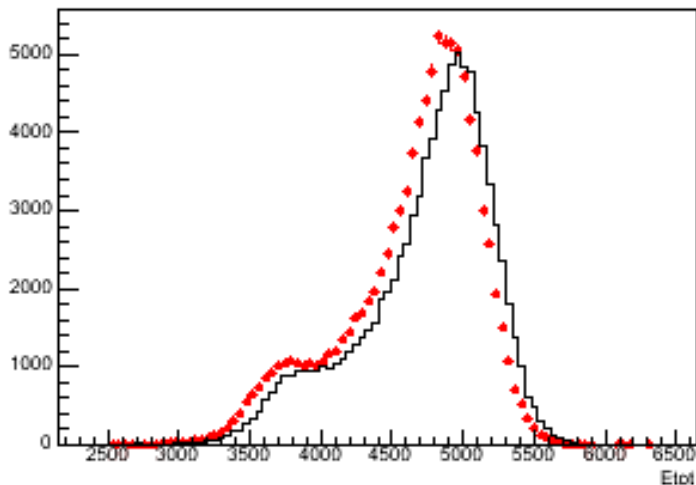
Xav {Tmax<100}



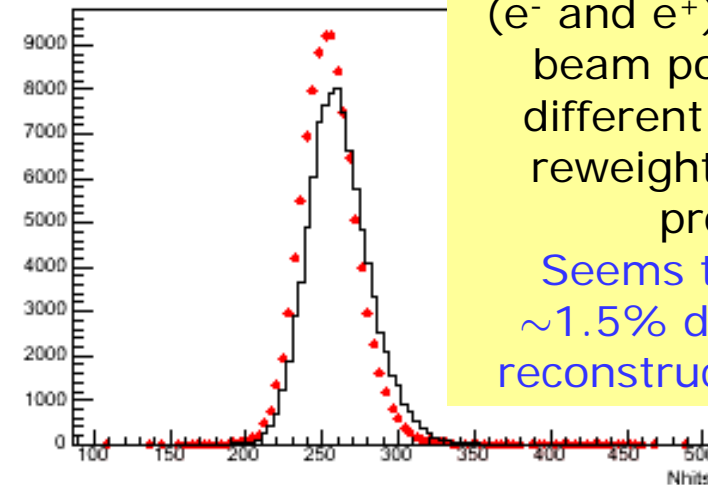
Yav {Tmax<100}



Etot {Tmax<100}



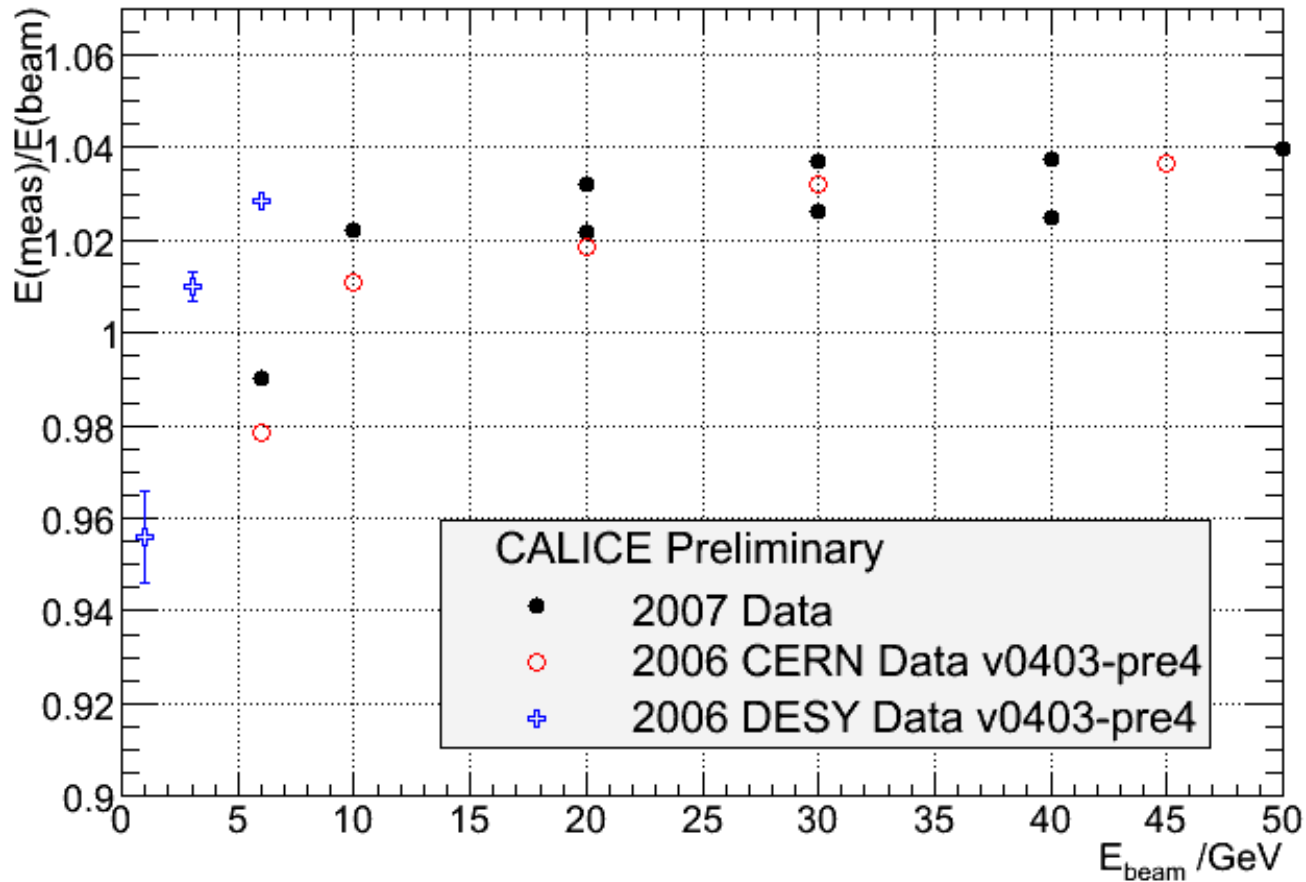
Nhits {Tmax<100}



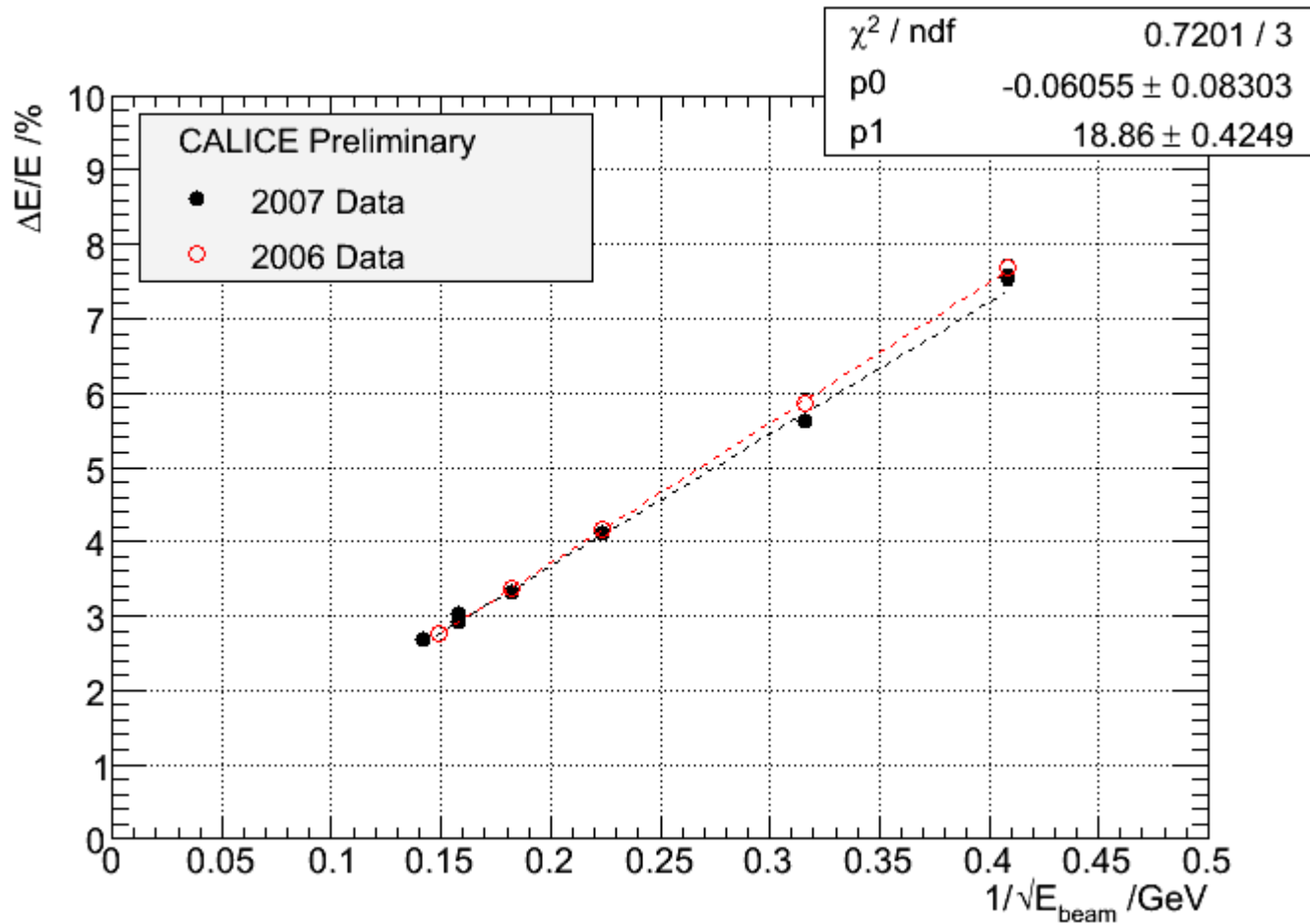
Take two 20 GeV runs (e⁻ and e⁺) with similar beam position, but different width, and reweight to ~same profile.

Seems to be clear
~1.5% difference in reconstructed energy.

Linearity 2006 *c.f.* 2007



Resolution 2006 *c.f.* 2007



Thoughts on systematic errors

- ❖ A major deficiency in what we've presented so far.
- ❖ Think what we might need to do, and whether software framework is adequate at present. Comments welcome.
 - ❖ Vary threshold cut.
 - ❖ Is the signal/noise cut in data reconstruction adequate?
 - ❖ Vary calibration constants within their statistical errors.
 - ❖ Reprocess using database? Or just smear? Or simulate using MC?
 - ❖ Vary alignment
 - ❖ Vary cuts for selecting good e^\pm events.
 - ❖ Compare runs at same nominal energy.
 - ❖ Vary fit procedure/range for extracting response/resolution
 - ❖ Different inter-wafer gap corrections?
 - ❖ Intrinsic beam energy uncertainty ($0.5\% \oplus 150 \text{ MeV}/p$) and spread (typically 0.5%).
- ❖ I think most of these could be done (or approximated well enough) using the existing reconstructed files.
- ❖ ...

...Continued...

- ❖ Would be good to add to the reconstructed files info about the beam (energy (nominal or “true”?), spread, position, angle, magnet currents?).
- ❖ Needs for the Monte Carlo?
 - ❖ Correct calorimeter geometry.
 - ❖ Including misalignments?
 - ❖ Dead cells (especially for 2007)
 - ❖ Signal-related crosstalk/pedestal shift, if we can understand the effect well enough.
 - ❖ More realistic simulation of the beam.
 - ❖ Spread of energy/position/angle; correlations between these.
 - ❖ Upstream material and showering.
 - ❖ Do we need a full beam line simulation?

Summary

- ❖ Do we go for publication of 2006 data soon, or wait till we understand 2007 data?
- ❖ Should we split up topics into smaller papers?
- ❖ Do we combine the DESY and CERN data?
- ❖ Big questions to resolve:
 - ❖ Understand/simulate the beam more correctly
 - ❖ Treatment of energy loss in inter-wafer gaps, in a way that is angle-independent.
 - ❖ Characterise/correct/simulate “square events” and other coherent signal-related effects, e.g. SIPS.
- ❖ Think realistically about systematics. We are not making precision SM measurements, but we do have to assess the reliability of anything we measure.